



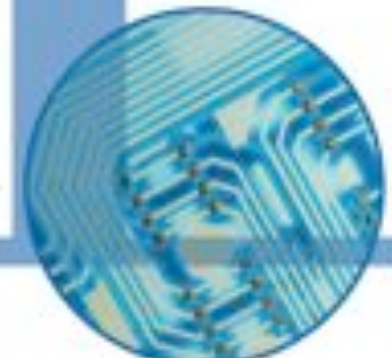
ADVANCED RADIO ASTRONOMY IN EUROPE

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Joint Institute for VLBI in
Europe
Dwingeloo, The Netherlands



ILIAS Annual meeting
Chambery, France, 26 February 2007

- Born on 5 May 1933
 - Karl Jansky, Bell Lab
 - $1 \text{ Jy} = 10^{-26} \text{ W}/(\text{m}^2\text{Hz})$
- The first non-visual window in the Universe (uv-, IR, X-ray, Gamma-ray and non-EM were to follow)
- Offers the deepest and sharpest view at the Universe

JENAM-2003 symposium

**Radio astronomy at 70:
from Karl Jansky
to microjansky**

Budapest, Hungary
27-30 August 2003




The New York Times

www.konkoly.hu/jenam03

Invited speakers:

W. Baan	A. Gurstein
R. Booth	H. Hirabayashi
A.G. de Bruyn	A. Konovalenko
B. Burke	M. Kramer
H. Butcher	J. Marcalde
D. Campbell	P. Mezger
J. Clayvil	I. Novikov
M. Cohen	J.M. Paredes
J. Condon	Yu. Pankov
T. Courvoisier	A. Readhead
E. Fomalont	R. Schilizzi
D. Galuzza	D. Schwartz
M. Garrett	A.R. Taylor
G. Giovannelli	G. Torricelli
S. Guilloteau	B. Wielebinski
	T. Wilson

Scientific/Organizing Committee:

R. Booth, A.G. de Bruyn, P. Diamond, S. Frey, L. Gurvits (chair), H. Hirabayashi, D.L. Jauncey, K. Kellermann, Yu. Pankov, S. Rawlings, R. Schilizzi, A. Szalay, P. Shaver, E. Valtajna, P.N. Wilkinson, A. Wolzycjan, J.A. Zensus

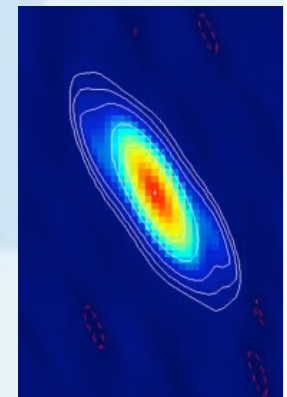
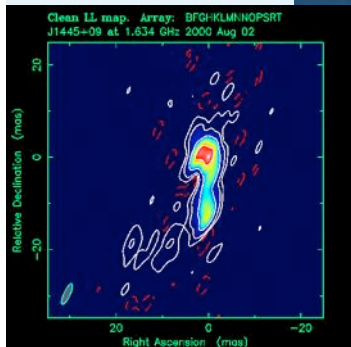
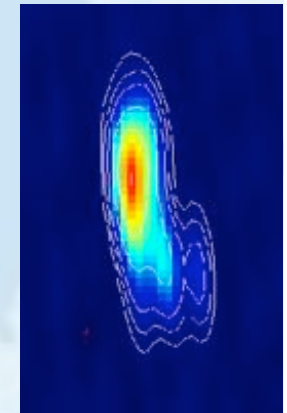
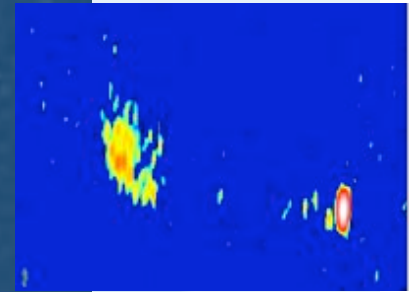
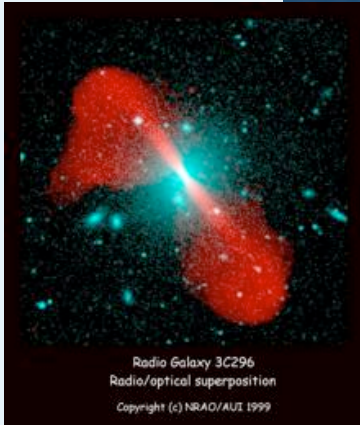
Local organizing Committee:

S. Frey (chair), L. Gurvits, C. Kiss, L. Mosoni

Sponsors:



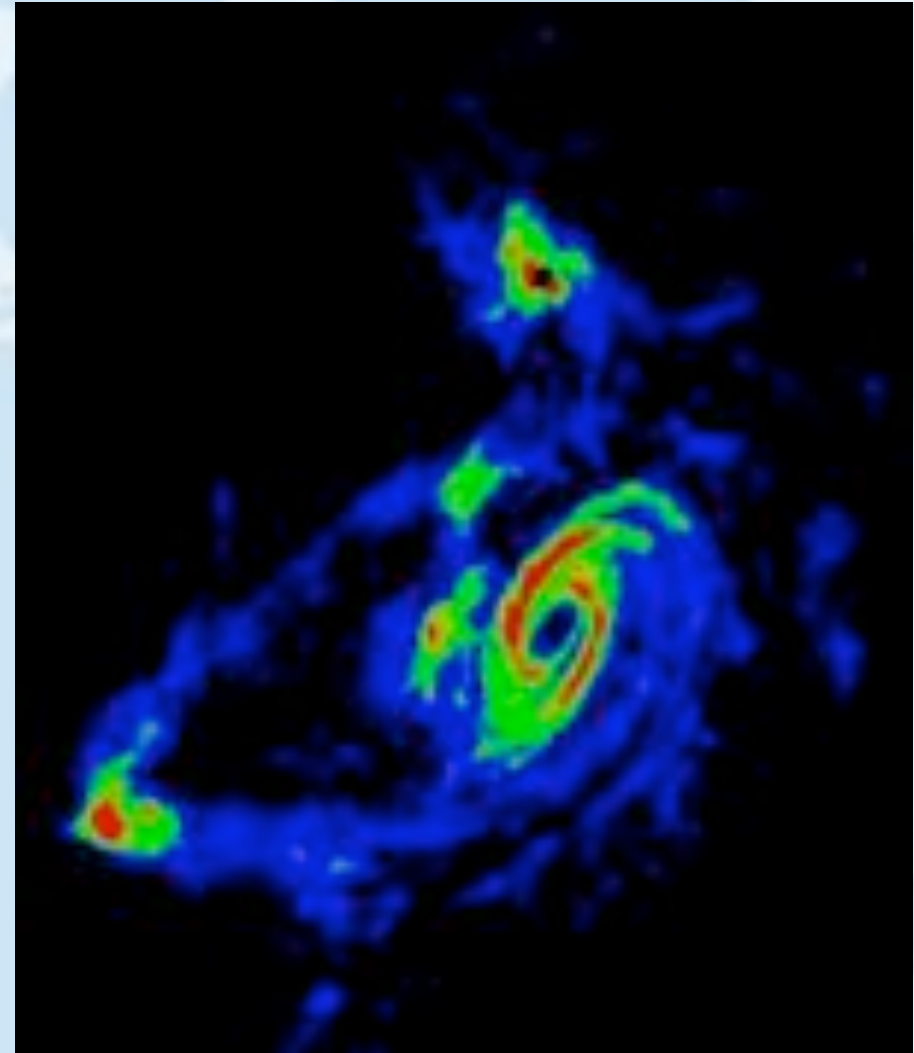
Sponsored by: European Astronomical Society, Hungarian Academy of Sciences, Eötvös Loránd University, Eötvös University, ASTRON, JIVE, Hungarian Scientific Research Fund and EU Infrastructure Cooperation Network RadioNET



A Universe of stars

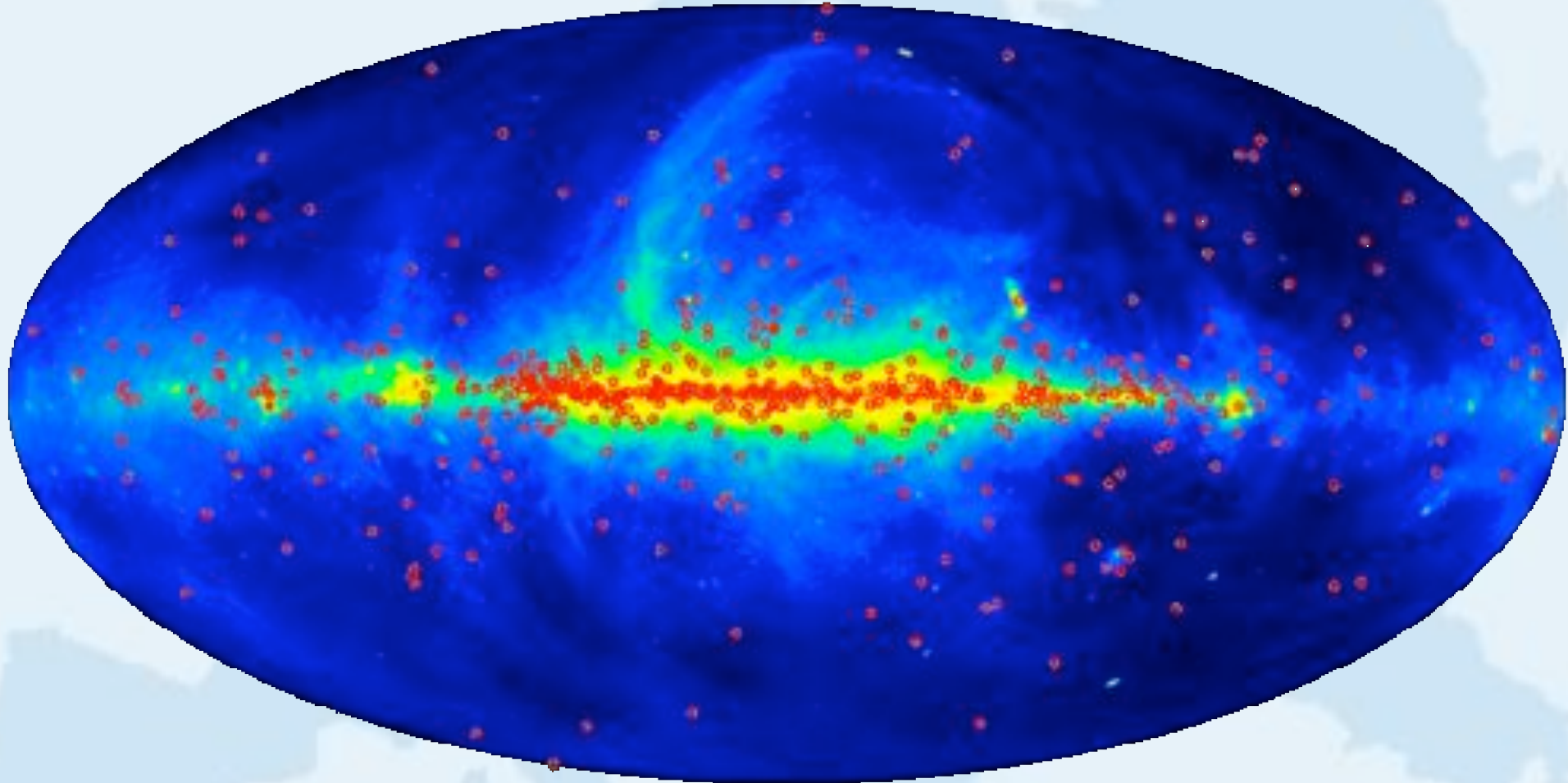


A Universe of hydrogen gas



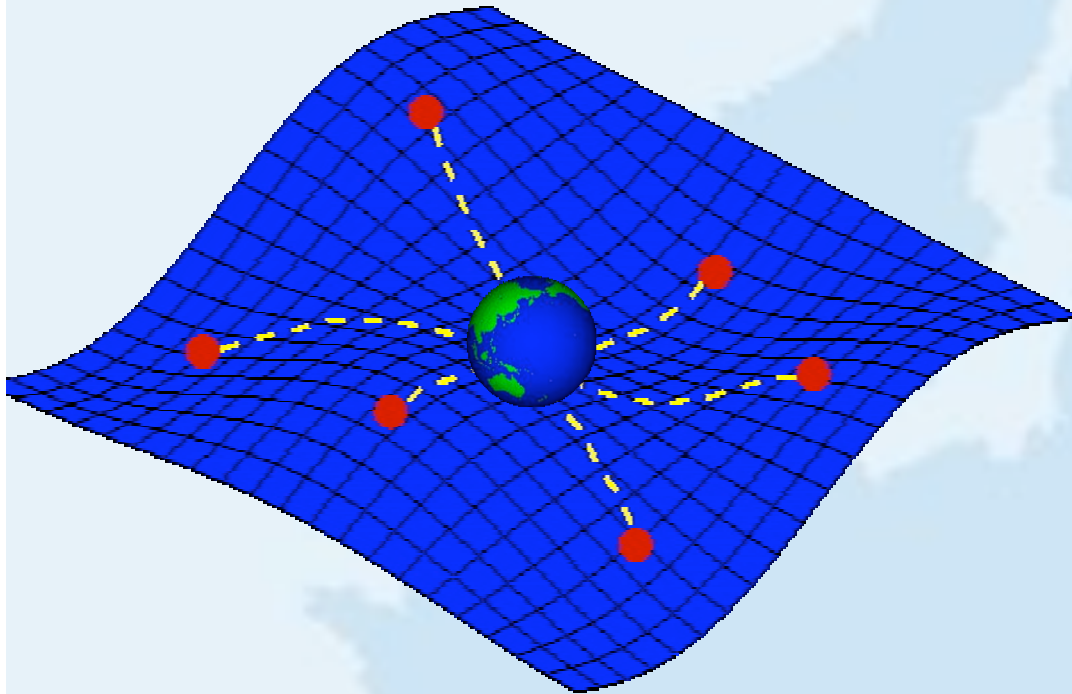
A very different view – (but HI 21 cm signal is weak!)

Pulsars – rotating neutron stars

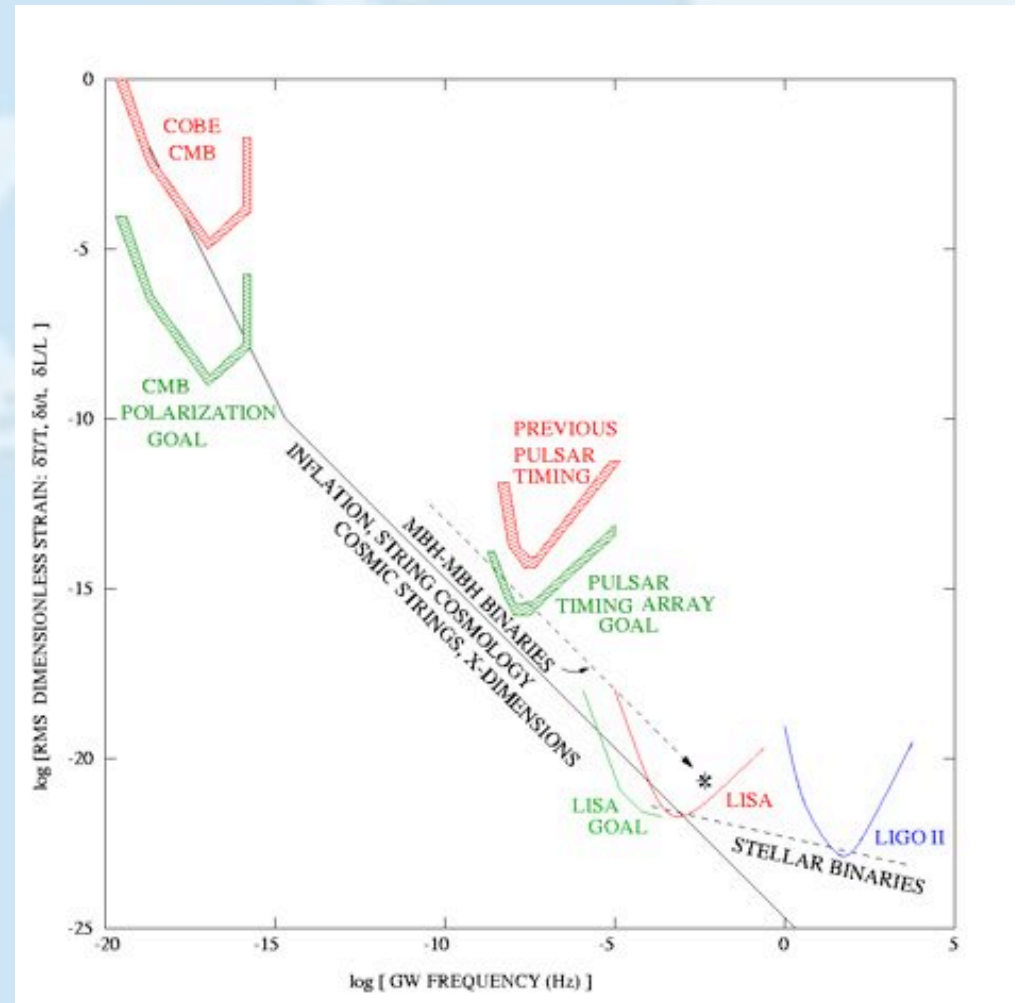


Born in the Galactic plane and then move away
(physics of initial “kick” not understood)

Timing a net of msec pulsars

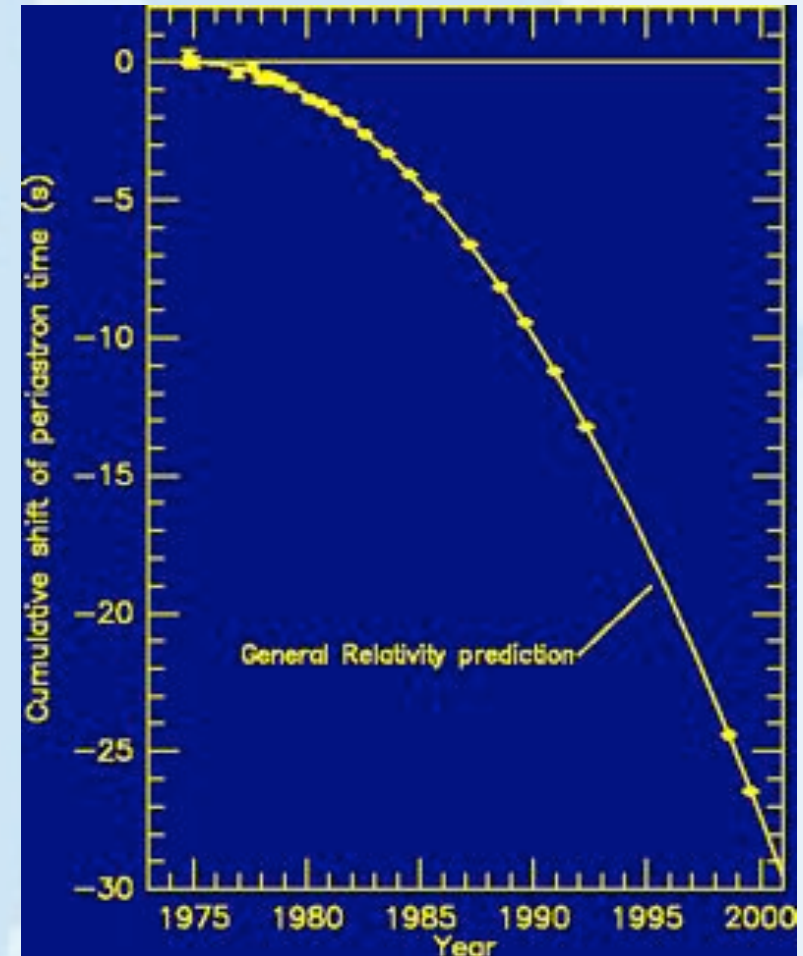
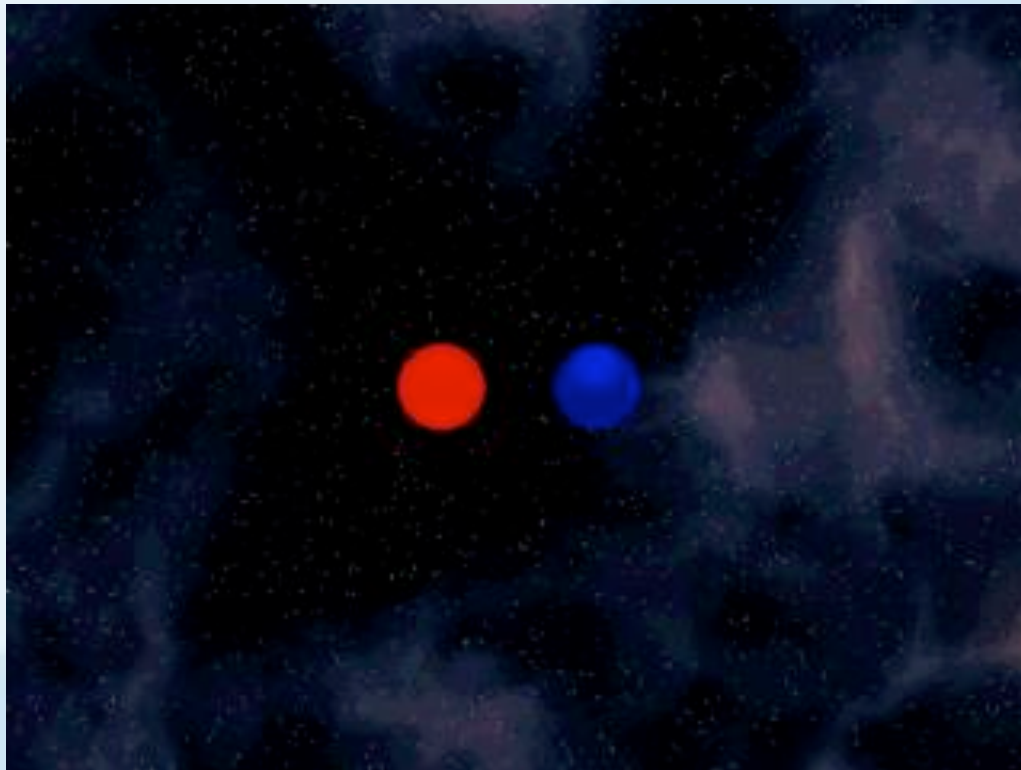


With SKA's timing precision look for spatial pattern in timing residuals!



- Pulsars=arms of huge gravitational wave detector
- Complementary to LIGO and LISA

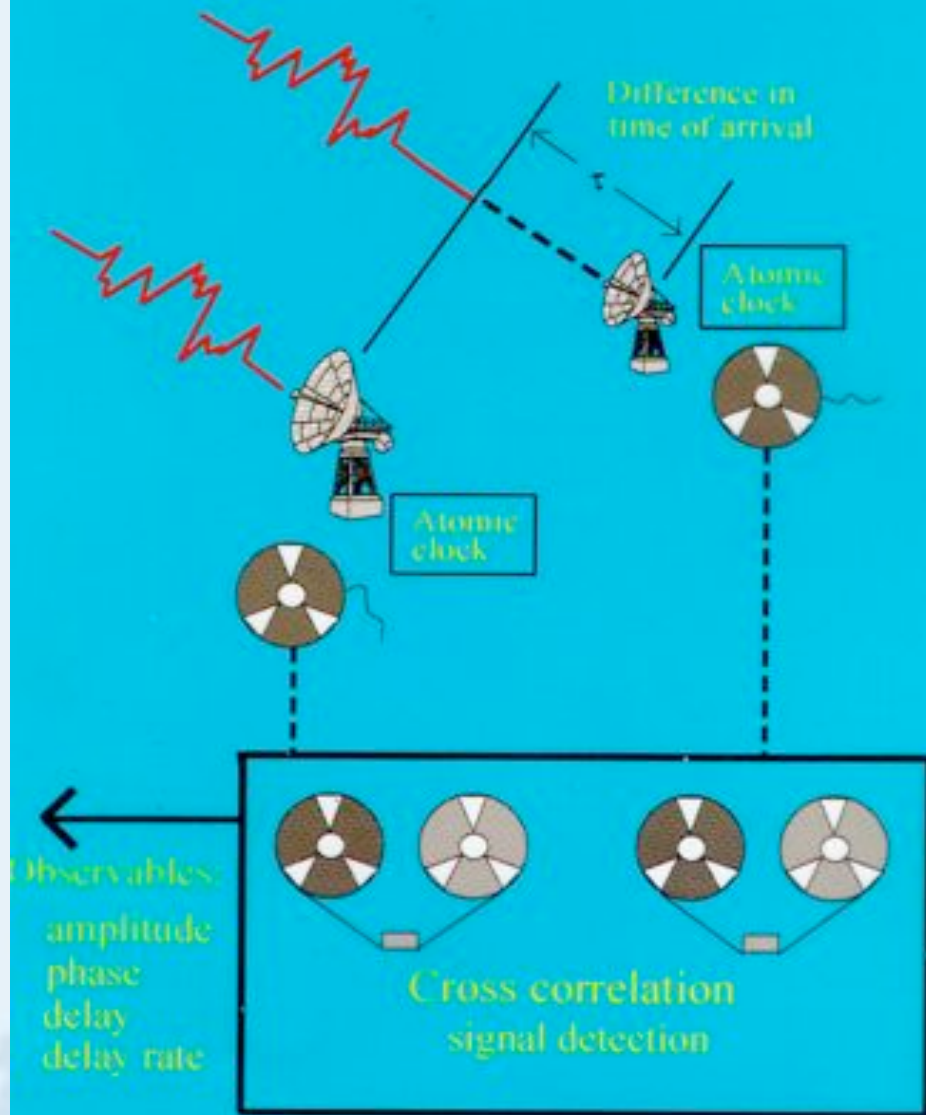
e.g. timing of B1913+16:



- Orbit shrinks every day by 1 cm!
- Confirmation of gravitational waves!



VLBI Configuration



Radio interferometry

← telescopes in different locations (countries, continents)

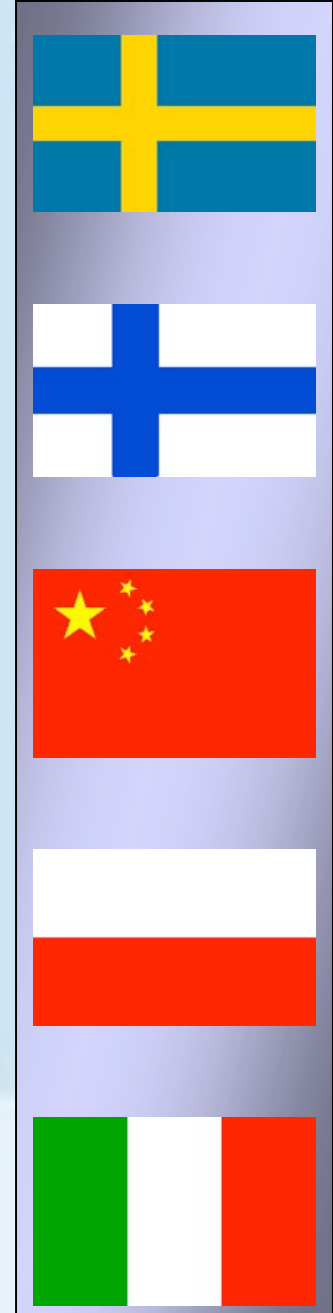
← data relayed or recorded on tape and transported to a central facility

← data processed (“correlated”)

Data rate: ~1 Gbps per RT;

Total: ~ 1-100 TB per exp.

The European VLBI Network



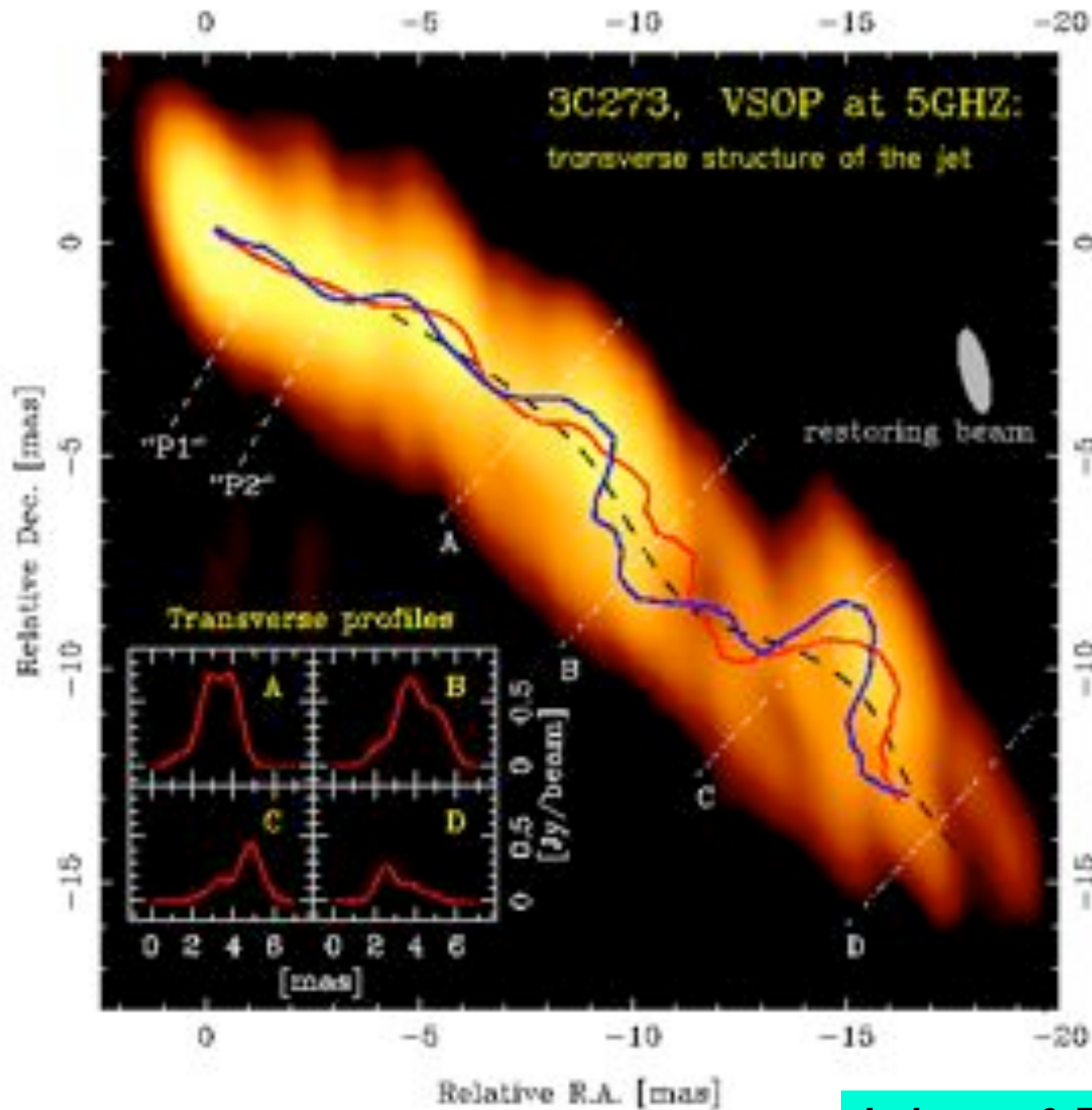
EVN data processor at JIVE (the Netherlands)



Arguably, the fastest computer in the world, BUT ...



**3C273:
the classical VLBI
target**



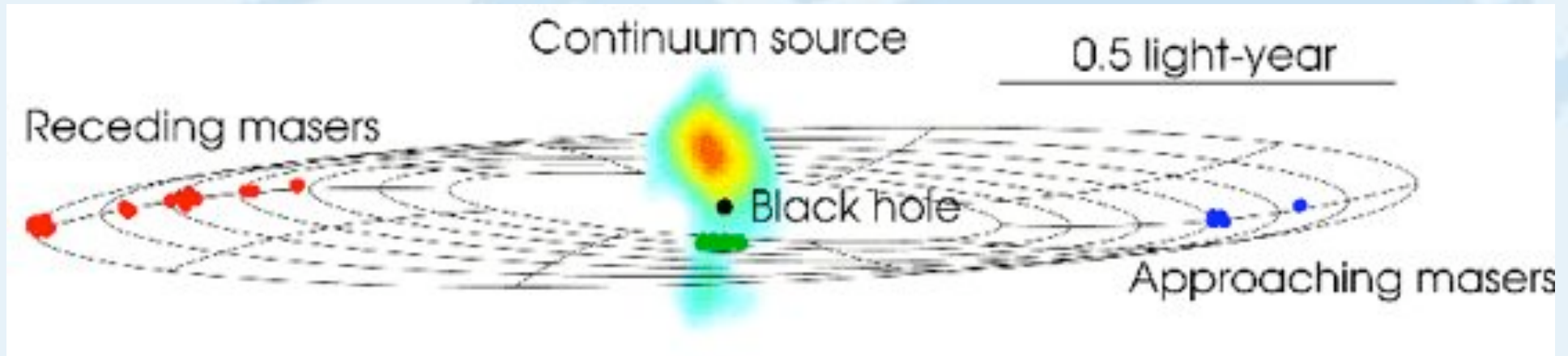
Lobanov & Zensus, 2000

NGC4258: the most convincing case for a super-massive black hole



(Miyoshi et al. 1995, Nat 373, 127; Herrnstein et al. 1998, ApJ 497, L69)

Rotating gas near the centre of the galaxy NGC4258, as traced by VLBI observations of the water vapor maser line at 22 GHz (1.35 cm)

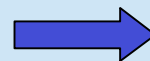


Doppler measurements

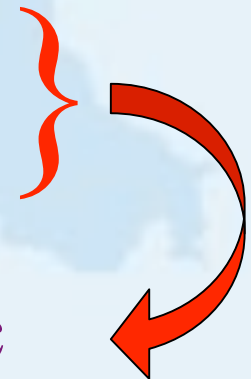


Radial velocities

VLBI measurements



(Angular) radii



$$V^2 / R \longrightarrow 3.6 \times 10^7 M_{\text{sun}} \text{ within } R = 0.13 \text{ pc}$$

VSOP/VLBA Pre-launch Survey of Extragalactic Radio Sources at 5 GHz (VLBApls)

E.B.Fomalont [1], S.Frey [2], Z.Paragi [2], L.I.Gurvits [3], W.K.Scott [4], A.R.Taylor [4], P.G.Edwards [5], H.Hirabayashi [5]

[1] National Radio Astronomy Observatory, Charlottesville, VA, USA

[2] University of Calgary, Alberta, Canada

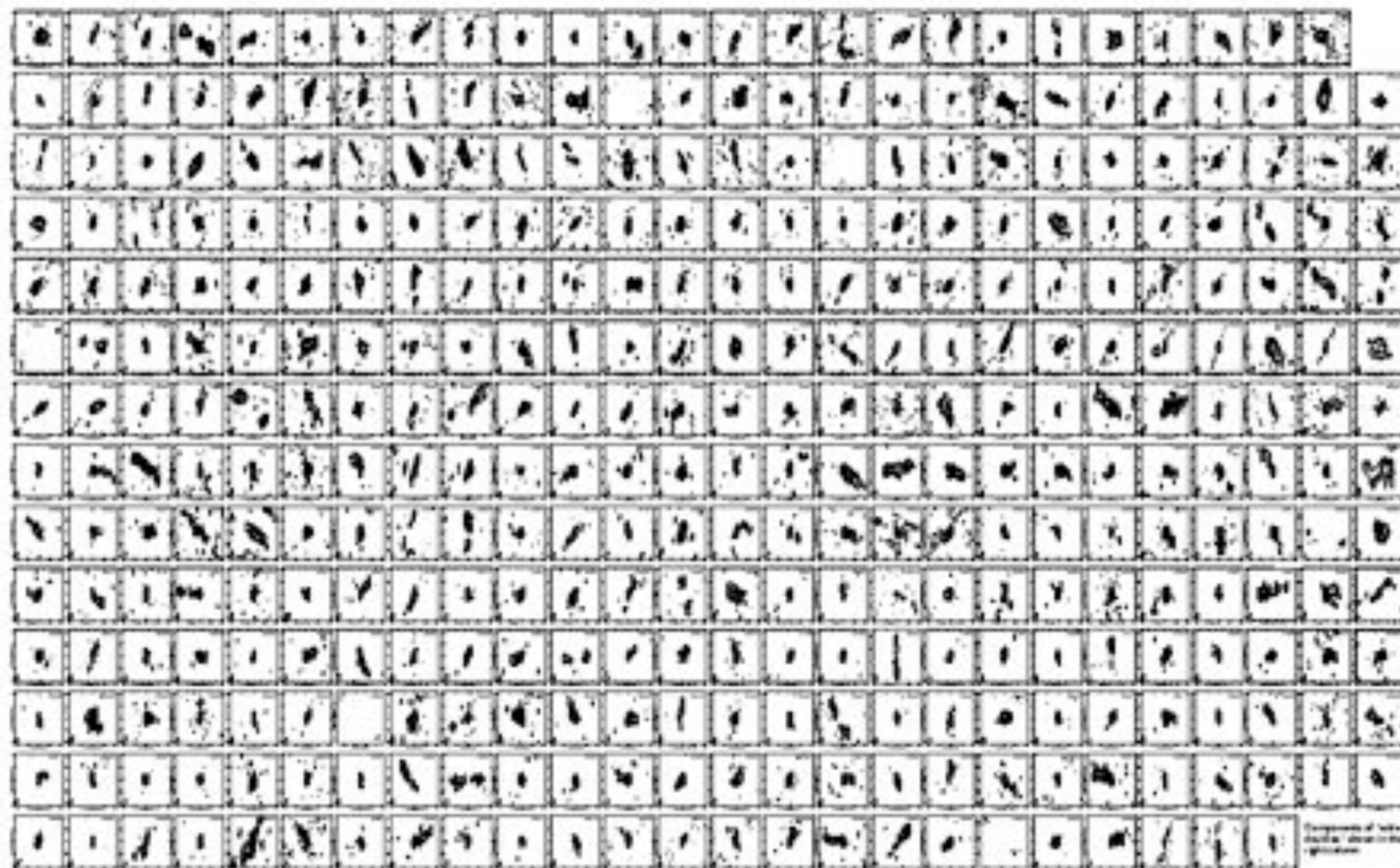
Observed 05-06 June 1995. Axes marked in milliarcsseconds.

[3] PDM Satellite Geodesy Observatory, Pécs, Hungary

[4] Institute of Space and Astronautical Science, Sagami-cho, Japan

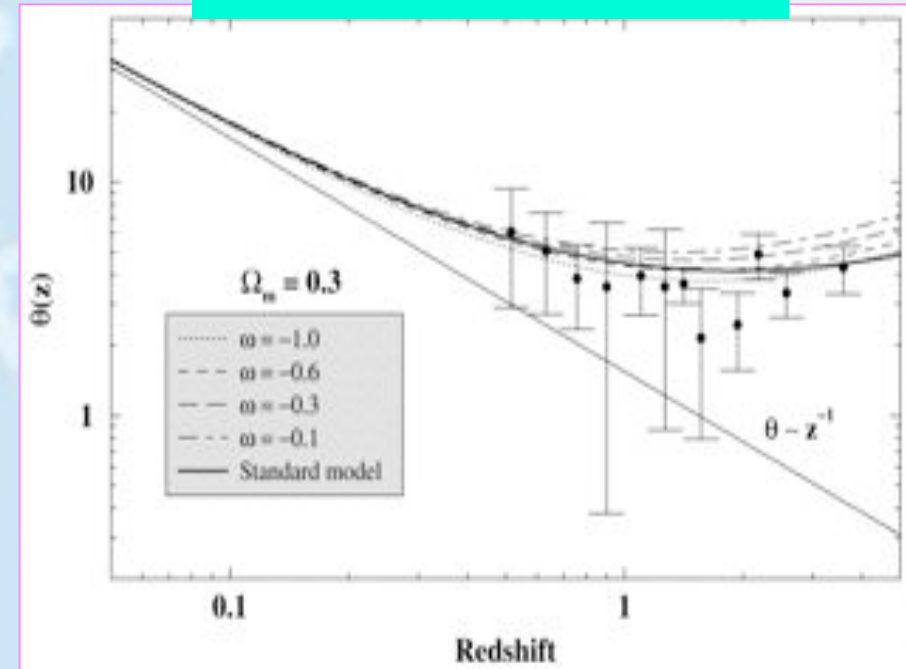
Astrophysical Journal Supplement, 2000

[5] Jodoh Institute for VLSI in Europe, Delft, The Netherlands

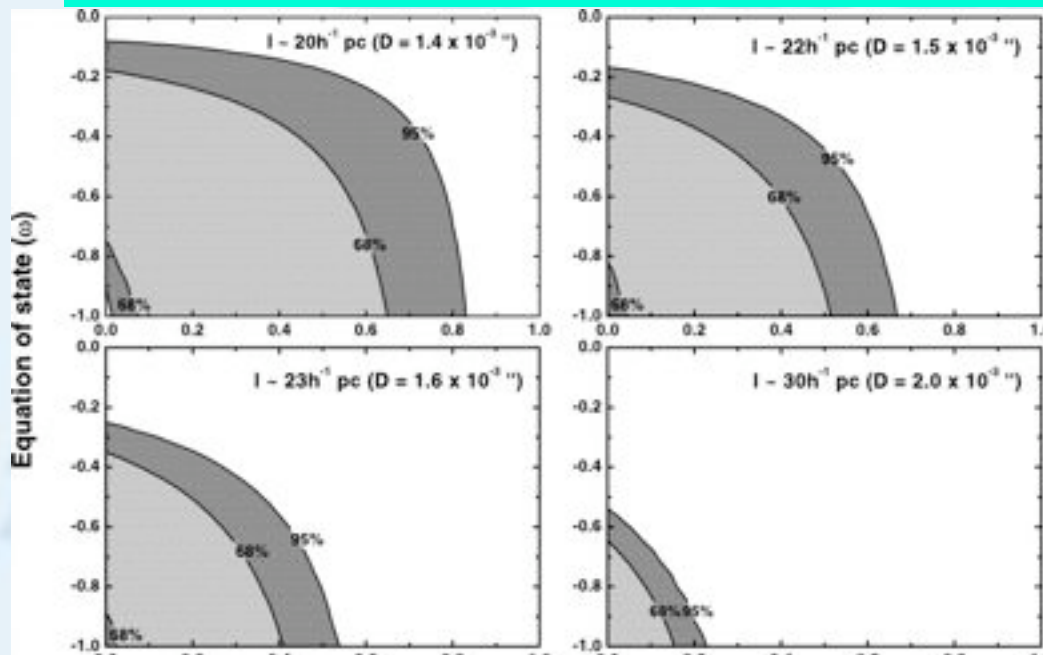


- FRW model driven by non-relativistic matter and a smooth “dark energy” component $p_x = \omega \rho_x$
- Best fit:

Lima & Alcaniz 2002



$\Omega_m \leq 0.62, \quad \omega \leq -0.2, \quad lh = 20 \text{ pc}$
 $\Omega_m \leq 0.17, \quad \omega \leq -0.65, \quad lh = 20 \text{ pc}$



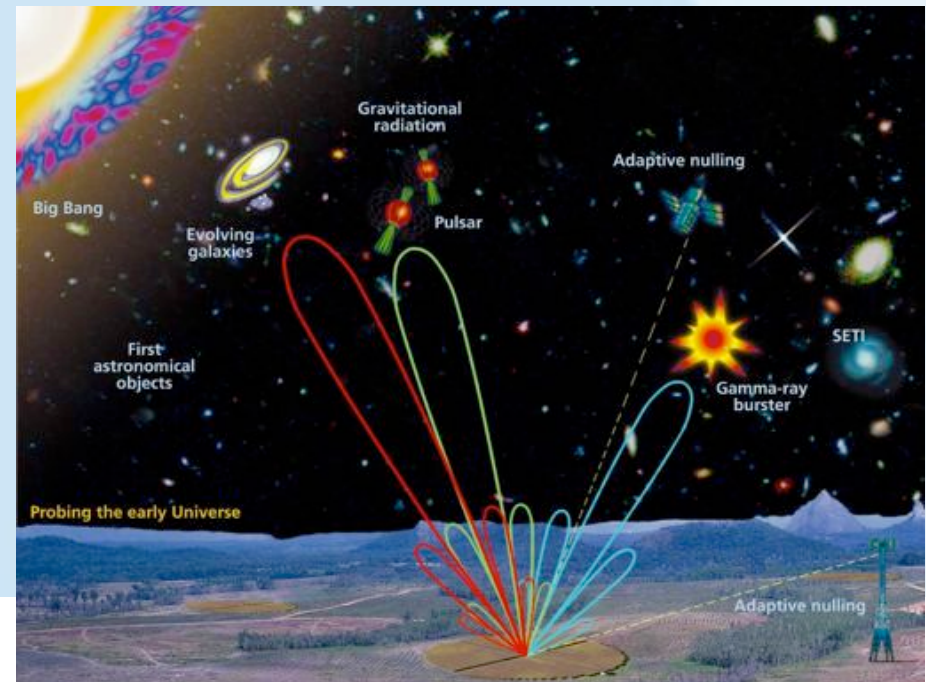
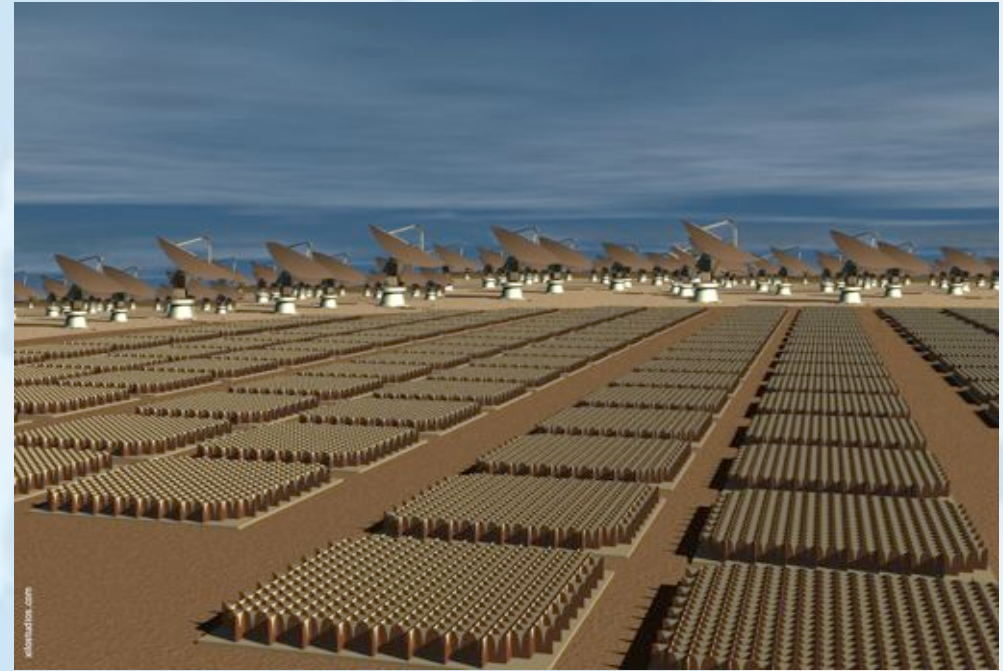
Conventional flat Λ CDM model ($\omega = -1$) with $\Omega_m = 0.2$ is the best fit.

Better statistics –

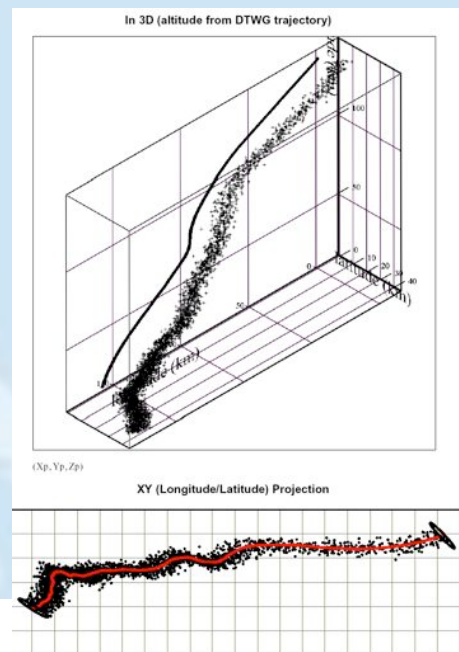
– more data needed!



- 100-fold increase in sensitivity
- Test-bed for new technologies and their applications beyond astronomy
- Natural bridge (in frequency domain) between LOFAR and ALMA
- Global project; to be constructed by 2020



- Next Generation VLBI correlator: pre-design study underway
- eVLBI EC FP6 project EXPReS funded; kick-off in 2006
- EC FP6 SKA Design Study commenced in 2005
- VLBI tracking of Huygens Titan Probe (Jan 2005), SMART-1 (May-Sep 2006)
- Active preparation for EC FP7 underway





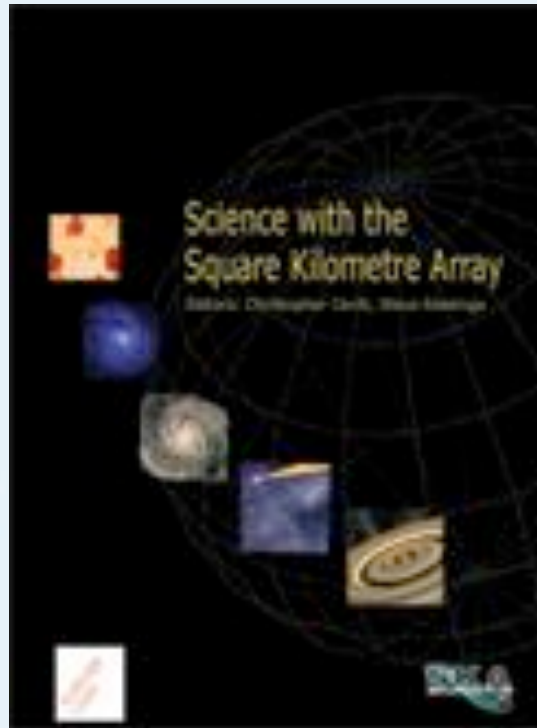
FP6 RadioNet at a glance

- Building-up on the strong European heritage in RA; descends from the 30-yr-old European VLBI Network
- Unifies the majority of RA institutes in Europe – 24 total
- Keeps RA developments on the edge-cutting level
- Total EC contribution 12.4 M€
- Provides synergy between various radio astronomy related technological developments, in particular:
 - FP6 SKA Design Study;
 - FP6 EXPRoS (e-VLBI, Info Society DG)
- Participates in the public outreach activities

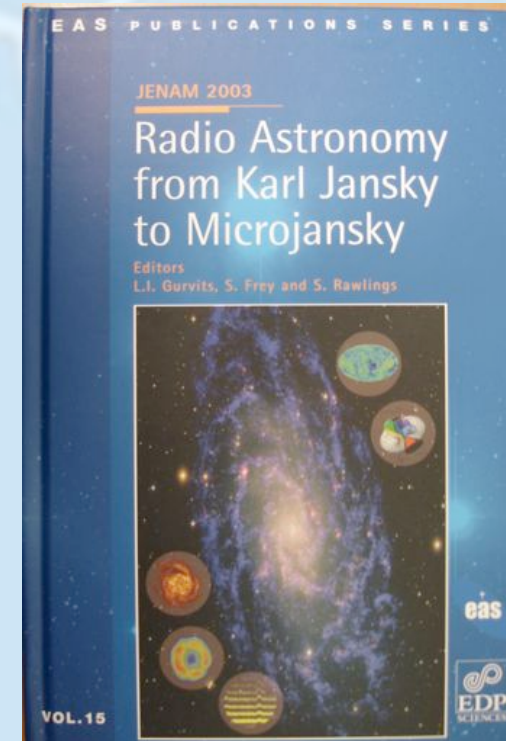
FP6 RadioNet at a glance (cntd)

- Three JRA's:
 - PHAROS: Aperture and focal array developments;
 - AMSTAR: mm radio astronomy LNA's and other instruments;
 - ALBUS: Advanced Long Baseline User Software
- Eight TNA facilities
- Eight Networks
- Governed by the Board (25 members, incl reps of sister I3's, ILIAS and OPTICON as at large members)
- Considers SKA as the focus of future developments (and collaborates closely with FP6 SKA Design Study)
- Stays in contact with AstroNet, OPTICON, ILIAS and EuroPlaNet on strategic planning issues

Recent publications



Science with the Square Kilometre Array
eds: C. Carilli, S. Rawlings,
New Astronomy Reviews,
Vol. 48, Elsevier, 2004
(see SKA booth, 2nd floor)



Radio Astronomy from Karl Jansky to Microjansky
eds: L.I. Gurvits, S. Frey, S. Rawlings
EAS Publ. Series, Vol. 15,
EDP Sciences, 2005
(see Astronomy and Astrophysics booth, 2nd floor)

More info on RadioNet : www.radionet-eu.org

RadioNet beyond FP6

- The strategy must be science-driven.
- Exploit new and strategic instruments owned and operated by Europe, e.g. e-MERLIN, e-EVN, PdB, upgraded single-dishes, LOFAR, ALMA and SKA
- Enunciate clear goal for the inclusion of a particular R&D area
- Ensure that we educate and train the next generation of astronomers and engineers:
 - Foster the leaders of the next decade.
- Think strategically and on European-scale, not nationally:
 - Relevant for RadioNet and also for new/upgraded infrastructure call

Decision-making

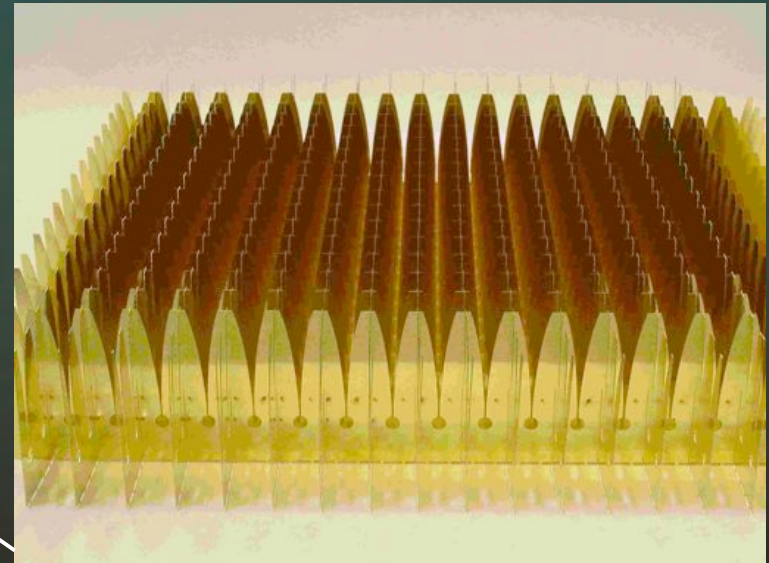
- Developing the decision-making process
 - TNA facilities must be world-class, relatively rare and offer a unique capability; they must have, or be able to demonstrate, a substantial European user base.
 - JRAs ideas to be developed further must fit within the strategic framework, they must have both a degree of relevance to existing facilities and can/should also demonstrate a role within future facilities.
 - NAs: Networking ideas should fit within the strategic framework of RadioNet.
- Should suggest areas of common interest for merging of ideas within RadioNet; possible interactions with other I3s?

RadioNet beyond FP6: JRA's

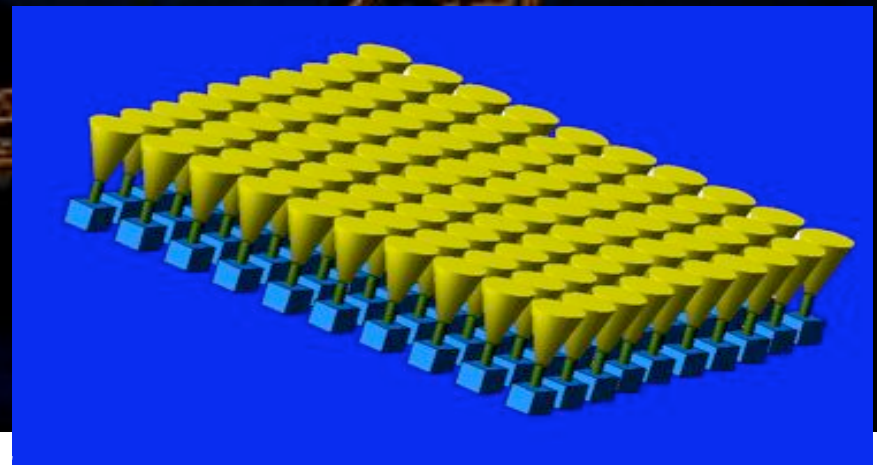
- JRAs – 5/6 major areas have emerged
 - Focal Plane Arrays (PHAROS+)
 - Very large format FPAs at mm/submm wavelengths and new methodologies at THz frequencies (AMSTAR+)
 - (User) software development (ALBUS+)
 - Digital systems: next generation VLBI correlator, pulsar timing, RFI monitoring/mitigation, SKA clock distribution, etc.
 - High-frequency and space science VLBI, astronomy from the Moon (prototyping)
 - SKA-related ideas (low power antennas, optical processing)

Focal Plane Arrays are science multipliers

Low-frequency (1.4-5 GHz) beam-forming arrays to maximise the potential of the EVN



- Mid-frequency (30-120 GHz) horn arrays to maximise the potential of large single dishes in Europe (e.g. Yebes, Effelsberg, SRT etc)
- High frequency (100-1000 GHz) horn arrays to maximise the potential of high-altitude single dishes in the era of ALMA



Networks

- Good case for continuing and expanding some, but not all, existing NAs:
 - Management
 - Synergy
 - Science Workshops
 - Engineering Forum
 - ALMA Forum
 - Spectrum monitoring

Networks

- See case for over-arching science workshop activity:
 - Coordinates workshops in different areas : general science themes, mm/sub-mm-related themes (separate in FP6); pulsar meetings; Suggestion on panchromatic workshops supporting SKA science case.
- See case for activity running schools & maybe science personnel exchanges:
 - m/dm/cm/mm/submm interferometry schools
 - Single-dish schools
 - YERAC
 - Solar physics schools
 - Spectrum management Schools
 - Training in best engineering practice

Networks

- Geonet – link geodesy to European astronomy activities
- LOFAR across Europe: planning, RFI, long-baseline calibration strategies...
- SKA non-astronomy applications
- QASP for E. European antennas
- Space VLBI – preparation for VSOP-2. Will now happen, so important we organise ourselves.
- ESKAC
- Policy / Industrial links

TNAs

- Define selection process – follow FP6 criteria?
 - Strong case for major existing facilities (but all should be re-examined)
 - Several major new instruments coming on-line soon – LOFAR, Yebes, SRT
 - Other large facilities : NRT, GMVA
 - Smaller, more focused facilities : INAF 32'ms, APEX (Swedish time), Nancay radioheliograph, AMI/VSA

- Can we define a strategy?
 - Prime – support of existing strategic facilities based on science goals. What are these goals? How do we define? Each facility has different but complementary science goals e.g. LOFAR vs PdB
 - Secondary – keep an eye on the future: SKA, ALMA, multi-waveband single-dish survey machines
 - How strong should SKA theme be within networking activities? They provide an avenue to support future facilities, I believe should be strong focus
- How do we strengthen links with other I3s and projects?
 - Inter-I3 JRA (e.g. Space Sci applications, with EuroPlaNet)
 - Joint ILIAS-RadioNet Science workshop on “dark” topics?



More on RadioNet

www.radionet-eu.org

